# Implementation of a Semi-implicit Orbit-Averaged Gyrokinetic Particle Code

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A semi-implicit orbit-averaged time-integration algorithm has been successfully implemented in a gyrokinetic particle simulation code for the study of self-consistent phenomena in a strongly magnetized plasma. The semi-implicit aspect of the integration scheme relaxes the timestep constraints required to ensure numerical stability. The orbit averaging is useful in reducing statistical noise and relaxes the statistical constraints for kinetic simulation. For appropriate applications, the semi-implicit orbit-averaged algorithm should be more efficient than are traditional particle-in-cell plasma simulation algorithms with explicit time-integration schemes. Both a linear numerical dispersion analysis and illustrative simulation examples are presented. © 1993

#### 1. INTRODUCTION

Kinetic phenomena in laboratory and space plasmas span many orders of magnitude making direct numerical solution of the fundamental equations computationally intensive (if not hopeless) in many circumstances [1]. This paper presents first results from the implementation of a semi-implicit orbit-averaged particle simulation algorithm that addresses the disparate timescale problem in the simulation of kinetic plasma phenomena. This study continues our earlier work [2] in which model algorithms and linear numerical dispersion analyses were presented that introduced both semi-implicit and semi-implicit orbitaveraged algorithms for self-consistent particle simulation of plasmas. Here we report our experience implementing a semi-implicit orbit-averaged time-integration scheme in an electrostatic gyrokinetic particle simulation algorithm with kinetic electrons and ions [3]. We have determined that modifications to the basic algorithm presented in [2] are required to suppress a slowly growing numerical error. A linear dispersion analysis of the revised algorithm and results from a few illustrative simulation examples are presented. The analysis and examples demonstrate that the semi-implicit orbit-averaged algorithm relaxes constraints

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on both the timestep required for stability of the highest frequency waves in the model and the statistical resolution. The combination of implicit time integration and orbit averaging leads to a more efficient self-consistent particle code for appropriate physics applications. However, the improvement in efficiency is sensitive to the details of the specific physics application.

The motivation for this work, that reported in [2], and the earlier work on direct implicit particle simulation [4] and orbit averaging [5, 6] is the desire to study time-dependent, collective kinetic phenomena in a plasma having a wide range of time scales so that the computational problem is quite stiff [1]. The class of phenomena motivating our specific choice of a gyrokinetic particle code as a testbed for a semi-implicit orbit-averaged time-integration scheme is low-frequency microturbulence in magnetically confined plasmas, e.g., drift-wave instabilities.

The gyrokinetic algorithm analytically removes the cyclotron timescale by a formal time averaging [3], while retaining the effects of the finite Larmor radius on the interaction of the particles with the self-consistent fields for frequencies much less than the cyclotron frequency. This assumption is compatible with a large class of drift-wave instabilities and low-frequency magnetohydrodynamic (MHD) and resistive modes that are believed to contribute significantly to the turbulent transport of energy and particles observed in tokamak experiments. Even with the use of the gyrokinetic formalism, the kinetic simulation of driftwave instabilities in tokamaks remains a stiff problem. The timescales of electrons transiting across a drift wave following the equilibrium magnetic field lines are many orders of magnitude shorter than the periods of the drift waves producing transport. Furthermore, an explicit time integration of the electrostatic equations with kinetic electrons is subject to a timestep constraint set by the highest frequency normal mode,  $\omega_h \Delta t < 1$ , where  $\omega_h \equiv (k_{\parallel}/k_{\perp})(m_i/m_e)^{1/2} \Omega_i$ ,  $\Omega_i = eB_0/m_i c$  is the ion cyclotron frequency,  $m_i/m_e$  is the mass ratio, and  $k_{\rm H}/k_{\perp}$  is the ratio of the parallel and perpendicular wavenumbers with respect to the equilibrium magnetic field;  $\omega_h$  is typically much larger than the driftwave frequencies. As explained in [2], the semi-implicit

orbit-averaged algorithm seeks to relax the timestep constraint set by  $\omega_h$  and reduce the statistical constraint on the electrons without forfeiting an accurate calculation of their trajectories.

The paper is organized as follows: In Section 2 we describe our implementation of the semi-implicit orbitaveraged gyrokinetic algorithm introduced in [2]. The linear dispersion analysis of this integration scheme in [2] indicated that exponentially growing modes can be stabilized at large timestep. However, trial simulations reveal numerically unstable zero-frequency modes that grow algebraically in time. A variant of this algorithm is then introduced, and a linear dispersion analysis shows that the stability boundary for exponentially growing modes is extended to a timestep that is significantly larger than the maximal stable timestep for the explicit time-integration scheme used in [3]. The effects of spatial filtering and interpolation onto a spatial grid modify the linear dispersion analysis. Simulation results for a few test cases using the revised algorithm are presented in Section 3. A simulation of an ion-temperature-gradient instability [7] is used as a convenient test problem with which to compare the simulation results of our algorithm and those from an explicit integration scheme. These results are in substantial agreement with those published in [7] for small timestep. In this example, finite-amplitude effects limit how large a timestep may be used and limit the usefulness and applicability of orbit averaging. Nevertheless, it is demonstrated that orbit averaging can lead to a substantial reduction of the number of particles. Additional simulations for a stable plasma are reported that illustrate the stability properties of the revised semi-implicit orbit-averaged algorithm and demonstrate both its noise-reduction characteristics and improved computational efficiency. Some concluding remarks are given in Section 4. The Appendix takes up a short discussion of how the algorithm can incorporate magnetic shear.

# 2. SEMI-IMPLICIT ORBIT-AVERAGED ALGORITHMS

In this section we shall begin by reviewing the algorithm introduced in [2] and describing both the results of a linear stability analysis and the failure of this algorithm because of a zero-frequency error that grows linearly in time. A revised algorithm is introduced that is designed to remove the secularly growing error. A linear dispersion analysis for the highest-frequency, cold-plasma normal mode is then presented. Some comments on how spatial grid effects and spatial filtering alter the linear dispersion relation are also given here.

### 2.1. First Algorithm

The first semi-implicit orbit-averaged algorithm was introduced in  $\lceil 2 \rceil$  for gyrokinetic simulation. The electrons

were modeled as a drift-kinetic species (zero-Larmor radius), while Larmor radius effects were retained in the gyrokinetic limit for the ions. Equations (17)-(21) in [2] describe the scheme in detail. The basic algorithm is composed of the following elements. In an electrostatic model, the electrons are advanced with a small timestep  $\Delta t_e$  from  $N \Delta t_i$  to  $(N+1) \Delta t_i$  using the electric field  $\mathbf{E}_N = -\nabla \phi_N$  in a static magnetic field. The electron charge and current densities  $n^e$  and  $J^e$  are accumulated on the spatial grid at each timestep from the particle data  $\{x^e, v^e\}$  and then timeaveraged to form  $\langle n^e \rangle$  and  $\langle \mathbf{J}^e \rangle_{N+1/2}$ . The ion gyrocenters are then advanced in a predictor step from  $N \Delta t_i$  to  $(N+1) \Delta t_i$  using the appropriately gyro-averaged electric field [3]. A predicted gyro-averaged ion number density  $\bar{n}^i$ is collected from the ion positions  $\{\tilde{\mathbf{x}}_{N+1}^i\}$  displaced by their Larmor radii. The self-consistent electric potential  $\phi_{N+1}$  is determined by the solution of the semi-implicit gyrokinetic Poisson equation,

$$-\left(\nabla^{2} + \frac{\omega_{pi}^{2}}{\Omega_{i}^{2}} \nabla_{\perp}^{2} g\right) (\tilde{\phi}_{N+1} - \phi_{N})$$

$$= 4\pi e (\bar{n}_{N+1}^{i} - \bar{n}_{N}^{i}) - 4\pi \Delta t_{i} \nabla \cdot \langle \mathbf{J}^{e} \rangle_{N+1/2}$$

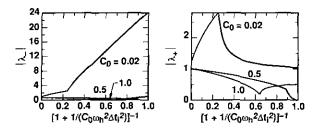
$$+ C_{0} \nabla_{||} [\bar{\omega}_{pe}^{2} \Delta t_{i}^{2} \nabla_{||} (\tilde{\phi}_{N+1} - \phi_{N})], \qquad (1)$$

where  $\bar{\omega}_{pe}^2 \equiv 4\pi (e^2/m_e) \langle n^e \rangle_{N+1/2} \approx 4\pi n_0 e^2/m_e$ ,  $\omega_{pi}^2 = 4\pi n_0 e^2/m_i$ ,  $g = b^{-1}[1 - I_0 \exp(-b)]$  in **k** space,  $b \equiv k_{\perp}^2 \bar{\rho}_i^2$ ,  $\bar{\rho}_i = (T_i/m_i)^{1/2}/\Omega_i$  is the ion thermal Larmor radius,  $\Omega_i$ is the ion cyclotron frequency, and  $C_0$  is a control parameter for the implicitness. Recall that the gyrokinetic ordering [2, 3] assumes that the density perturbations considered, although finite, are small in amplitude, and the inhomogeneities of the unperturbed plasma are weak. In consequence, the plasma parameters appearing in the ion polarization and the semi-implicit electron susceptibility are homogeneous and constant. The ions are then advanced in a corrector step from  $N \Delta t_i$  to  $(N+1) \Delta t_i$  using  $\tilde{\mathbf{E}}_{N+1} =$  $-\nabla \tilde{\phi}_{N+1}$  and  $\mathbf{E}_N$ . The corrected value of  $\bar{n}_{N+1}^i$  is then calculated and used in Eq. (1) to obtain  $\phi_{N+1}$ . The timestep advance to  $(N+1) \Delta t_i$  is completed by using an electric field  $\mathbf{E} = \mathbf{E}_N + 2C_0(\mathbf{E}_{N+1} - \mathbf{E}_N)$ , where  $\mathbf{E}_{N+1} = -\nabla \phi_{N+1}$  to correct the advance of the electrons from  $N \Delta t_i$  to  $(N+1) \Delta t_i$ .

A linear stability analysis was performed in [2] to obtain the model dispersion relation in the limit of a cold, nondrifting plasma and with spatial grid and interpolation effects neglected,

$$(1 + C_0 \omega_h^2 \Delta t_i^2)(\lambda - 1)^2 + \omega_h^2 \Delta t_i^2 [(2C_0 + 1/2)(\lambda - 1) + 1] = 0,$$
 (2)

where  $\lambda = \exp(-i\omega \Delta t_i)$ ,  $\omega_h^2 \equiv (k_{\parallel}^2/k_{\perp}^2)(m_i/m_e) \Omega_i^2$ , and  $k_{\parallel}^2 \ll k_{\perp}^2$ . The parallel and perpendicular subscripts indicate



**FIG. 1.** The modulus of the amplification factor  $|\lambda|$  as a function of  $\omega_B^2 \Delta t_i^2$  for  $C_0 = 0.02, 0.5, 1.0$  in the first algorithm.

the orientation of the wavenumber components relative to the magnetic field. For  $C_0 > \frac{1}{4}$ , there are damped solutions  $(|\lambda| < 1)$  of Eq. (2): for  $\omega_h^2 \Delta t_i^2 \gg 1$ ,  $\lambda = \{1 - 1/(2C_0), -1 + O(1/\omega_h^2 \Delta t_i^2)\}$ ; for  $\omega_h^2 \Delta t_i^2 \ll 1$  there are damped solutions with  $\text{Re}(\omega) \approx \pm \omega_h$ . In Fig. 1 we plot the modulus of the amplification factor for the two roots of Eq. (2). Thus, for  $C_0 > \frac{1}{4}$  there are no exponentially growing modes. However, perturbations growing with a power-law dependence on time are not excluded.

In our unsuccessful attempts to implement this algorithm, we discovered several difficulties. The first difficulty was that Eq. (1) treats the ion and electron source terms in Eq. (1) on an unequal footing. The electron current was accumulated and then its divergence calculated on the grid, while an ion charge density was calculated directly. To put the electrons and ions on a more equal footing with respect to finitedifferencing, which might be important in resolving small deviations from quasi-neutrality, we made the replacement  $4\pi e(\bar{n}_{N+1}^{i} - \bar{n}_{N}^{i}) = -4\pi \Delta t_{i} \nabla \cdot \tilde{\mathbf{J}}_{N+1}^{i}$  in Eq. (1), where  $\tilde{\mathbf{J}}_{N+1}^{i}$ is the gyro-averaged ion current density at  $(N+1) \Delta t_i$  on the ion predictor step and is determined from  $\{\tilde{\mathbf{x}}, \tilde{\mathbf{v}}\}_{N+1}^{i}$ . On the ion corrector step,  $J_{N+1}^{i}$  was accumulated from  $\{\mathbf{x}, \mathbf{v}\}_{N+1}^{i}$ . This modification leaves the dispersion relation derived in Eq. (2) unchanged. However, a second difficulty remains and manifests itself in zero-frequency errors in the potential and perturbed plasma displacements that grow approximately linearly in time (Fig. 2). Although we were

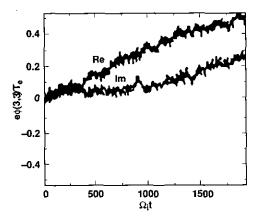


FIG. 2. The real and imaginary parts of the  $(k_x, k_y) = (3, 3)$  Fourier component of  $e\phi/T_e$  as a function of time.

able to verify the algorithm's stability properties with respect to exponentially growing modes for large values of  $\omega_h^2 \Delta t_i^2$  when  $C_0 > \frac{1}{4}$ , or instability when  $C_0 < \frac{1}{4}$ , the presence of the secularly growing error frustrated successful simulation of physical instabilities like the ion-temperature-gradient (ITG) mode [7]. We believe that the secularly growing errors derive from grid errors and statistical effects in the ion and electron source terms on the right side of Eq. (1) that linearly accrue in  $\phi$  at each timestep on account of the time integration of Eq. (1). In k space, Eq. (1) has the form  $\phi_{N+1} = \phi_N + L_k S_k$ , where  $L_k$  is a linear operator and  $S_k$  is a source. By replacing Eq. (1), with a semi-implicit Poisson equation that uses the ion charge density and the orbit-averaged electron charge density, we were able to remove the secularly growing error.

## 2.2. Revised Algorithm

We have revised our semi-implicit orbit-averaged algorithm to remove the secularly growing errors as follows. The interleaving of the ion and electron advances is diagrammed in Fig. 3. The electrons are advanced in a predictor step from  $(N+1/2) \Delta t_i$  to  $(N+3/2) \Delta t_i$  using a small timestep  $\Delta t_e$  and the electric field  $\mathbf{E}_N$ . The orbit-averaged electron number density  $\langle n^e \rangle_{N+1}$  is accumulated by averaging the electron densities calculated at each electron timestep over this time interval. The ions are advanced in a predictor step from  $N \Delta t_i$  to  $(N+1) \Delta t_i$  using the larger timestep  $\Delta t_i$  and the electric field  $\mathbf{E}_{N+1}$  in a conventional way [2, 3]. A gyroaveraged ion number density  $\tilde{n}_{N+1}^i$  is accumulated at  $(N+1) \Delta t_i$ . The revised semi-implicit gyrokinetic Poisson equation is solved following the ion advance,

$$-\left(\nabla^{2} + \frac{\omega_{pi}^{2}}{\Omega_{i}^{2}}\nabla_{\perp}^{2}g\right)\tilde{\phi}_{N+1} = 4\pi e(\tilde{n}_{N+1}^{i} - \langle n^{e}\rangle_{N+1}) + \nabla \cdot (C_{0}\chi_{e} \cdot \nabla(\tilde{\phi}_{N+1} - \phi_{N})),$$
(3)

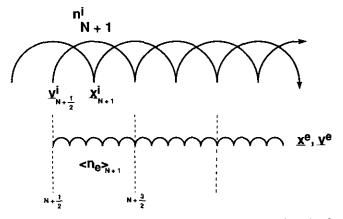


FIG. 3. Interleaving of advancing ions and electrons and solving for he self-consistent electric field in the revised orbit-averaged algorithm.

where  $\chi_e$  is the semi-implicit susceptibility tensor and is defined in the following.

The susceptibility is derived by relating the linearized drift-kinetic electron response to the electric field. For a slab configuration with  $\mathbf{B} = B_0 \hat{\mathbf{z}} + B_y \hat{\mathbf{y}}$ ,  $|B_y/B_0| \le 1$ , and equilibrium density gradient in the x direction, both the parallel electron response and the perpendicular  $\mathbf{E} \times \mathbf{B}$  drift lead to charge density perturbations. Hence, Eq. (3) in the drift-kinetic limit for the electrons becomes

$$-\left(\nabla^{2} + \frac{\omega_{pi}^{2}}{\Omega_{i}^{2}}\nabla_{\perp}^{2}g\right)\widetilde{\phi}_{N+1} - C_{0}\left(\nabla_{\parallel}\widetilde{\omega}_{pe}^{2}\Delta t_{i}^{2}\nabla_{\parallel}\right)$$

$$+2\frac{\widetilde{\omega}_{pe}^{2}\Delta t_{i}^{2}}{\Omega_{e}\Delta t_{i}}\frac{1}{L_{n}}\frac{\partial}{\partial y}\left(\widetilde{\phi}_{N+1} - \phi_{N}\right)$$

$$= 4\pi e(\overline{n}_{N+1}^{i} - \langle n^{e} \rangle_{N+1}). \tag{4}$$

The two semi-implicit terms contribute to relaxing the  $\omega_h \Delta t_i$  stability constraint and to compensating for the lack of centering in the electron charge density response which impacts the temporal accuracy. The validity criterion for the linearization leading to the semi-implicit susceptibility is obtained from the straightforward extension of the corresponding condition derived for the direct implicit method [4, 8] to a magnetized plasma,  $|k_{\parallel}^2(e/m_e) \phi \Delta t_i^2| < 1$ . This condition is a statement that the parallel displacement produced by the acceleration due to the electric field during a timestep must be smaller than a relevant wavelength to be resolved. This is a timestep constraint on the integration of the electron trajectory that must be satisfied whether or not the integration scheme is implicit and does not limit the nonlinearity of the electron response.

The ions are advanced in a corrector step from  $N \Delta t_i$  to  $(N+1) \Delta t_i$  using  $\tilde{\mathbf{E}}_{N+1} = -\nabla \tilde{\phi}_{N+1}$  and  $\mathbf{E}_N$  as described in [2, 3]. The corrected ion positions  $\{\mathbf{x}_{N+1}^i\}$  are used to accumulate  $\bar{n}_{N+1}^i$ , and Eq. (4) is solved to determine the corrected value of  $\phi_{N+1}$  from which  $\mathbf{E}_{N+1} = -\nabla \phi_{N+1}$ . Next the electron trajectories from  $(N+1/2) \Delta t_i$  to  $(N+3/2) \Delta t_i$  are recomputed in a corrector step again using a small timestep  $\Delta t_e$ . The electrons are advanced with an electric field  $\mathbf{E} = C_0 \tilde{\mathbf{E}}_{N+1} + (1-C_0) \mathbf{E}_N$  for this corrector step. This completes the computational cycle.

The highest frequency normal mode supported by these equations in a cold uniform plasma has the frequency  $\omega_h$  defined in Section 1. The linear dispersion relation including  $\Delta t$  effects is obtained straightforwardly using the methods of Refs. [1, 2, 5, 6]. For the  $\omega_h$  mode, the parallel electron response dominates the right side of Eq. (4); and  $(\omega_{pi}^2/\Omega_i^2) \nabla_{\perp}^2 g \rightarrow (\omega_{pi}^2/\Omega_i^2) \nabla_{\perp}^2 \gg \nabla^2$  for a cold, high-density plasma. For a small-amplitude, linear charge density perturbation, the right side of Eq. (4) becomes

$$4\pi e(\hat{n}_{N+1}^{i} - \langle n^{e} \rangle_{N+1}) \approx 4\pi e \nabla_{||} \left[ n_{0} \sum_{j=0}^{N_{*}} x_{||j}^{e} / (N_{*} + 1) \right]$$

$$= 4\pi e \nabla_{||} \left[ n_{0} \left( x_{||N+1/2}^{e} + \frac{\Delta t_{i}}{2} v_{||N+1/2}^{e} + \left( \frac{-e}{m_{e}} \right) \frac{\Delta t_{i}^{2}}{6} E_{||N} \right) \right], \tag{5}$$

where  $N_* = \Delta t_i/\Delta t_e$  (we assume that  $N_* \ge 1$ ). Note that Eq. (5) corrects a typographical error in Eq. (7) of [6].

The linearized electron trajectory can be reduced to

$$x_{||N+3/2}^{e} - x_{||N+1/2}^{e} = \frac{\Delta t_{i}}{2} \left( v_{||N+1/2}^{e} + v_{||N+3/2}^{e} \right)$$
 (6a)

$$v_{||N+3/2}^{e} - v_{||N+1/2}^{e} = -\frac{e}{m_{e}} \Delta t_{i} E_{||}^{*},$$
 (6b)

where  $\mathbf{E}^* = \mathbf{E}_N$  on the predictor advance and  $\mathbf{E}^* = C_0 \mathbf{E}_{N+1} + (1-C_0) \mathbf{E}_N$  on the corrector advance;  $\phi_{N+1}$  is related by Eqs. (4) and (5) to  $x^e_{\parallel N+1/2}$ ,  $v^e_{\parallel N+1/2}$ , and  $\phi_N$ . Using the corrector version of Eqs. (6a) and (6b), we can relate the linearized values of  $x^e_{\parallel}$  and  $v^e_{\parallel}$  directly to  $E_{\parallel N}$  and  $E_{\parallel N+1}$ . We then introduce the amplification factor  $\lambda = \exp(-i\omega \Delta t_i)$  and Fourier transform the spatial dependence to obtain

$$\tilde{x}^{e} = -\frac{\Delta t_{i}^{2}}{2} \frac{e}{m_{e}} \frac{\lambda^{1/2} (\lambda + 1)}{(\lambda - 1)^{2}} \left[ C_{0} + \frac{(1 - C_{0})}{\lambda} \right] i k_{||} \tilde{\phi}$$
 (7a)

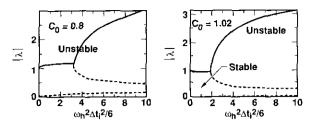
$$\tilde{v}^e = -\Delta t_i \frac{e}{m_e} \frac{\lambda^{1/2}}{(\lambda - 1)} \left[ C_0 + \frac{(1 - C_0)}{\lambda} \right] i k_{11} \tilde{\phi}, \qquad (7b)$$

where  $\tilde{x}^e$ ,  $\tilde{v}^e$ , and  $\tilde{\phi}$  are Fourier amplitudes. Equations (4) and (5) are also Fourier analyzed to remove the spatial dependence, and we substitute  $\tilde{E}_{\parallel} = -ik_{\parallel}\tilde{\phi}$ , and  $\tilde{x}^e$  and  $\tilde{v}^e$  from Eqs. (7a) and (7b) to obtain the dispersion relation

$$\lambda(\lambda - 1)^2 + \alpha(\lambda^2 + 4\lambda + 1)(C_0\lambda + 1 - C_0) = 0,$$
 (8)

where  $\alpha \equiv \omega_h^2 \Delta t_i^2 / 6$ .

There are three roots of Eq. (8) in general. For  $\alpha \gg 1$ , there is one root  $\lambda = (C_0 - 1)/C_0$  that is stable for  $C_0 \gg \frac{1}{2}$  and two unstable roots  $\lambda = -2 \pm \sqrt{3}$ . For  $C_0 = 1$  there is one stable root  $\lambda = 0$  and two additional roots that are stable if  $\alpha \leqslant 2$ , i.e.,  $\omega_h^2 \Delta t_i^2 \leqslant 12$ . The general solution of Eq. (8) indicates that there is at least one unstable root for  $C_0 < 1$  or  $\omega_h^2 \Delta t_i^2 > 12$ . As  $C_0$  is increased above  $C_0 = 1$ , the instability threshold on the timestep decreases below  $\omega_h^2 \Delta t_i^2 = 12$ . Sample solutions for  $|\lambda|$  as a function of  $\alpha$  are plotted in Fig. 4. We conclude that the revised semi-implicit orbit-averaged algorithm relaxes the timestep constraint from  $\omega_{11}^2 \Delta t_i^2 \leqslant O(1)$  to  $\omega_h^2 \Delta t_i^2 \leqslant 12$ . We note that in contrast



**FIG. 4.** The modulus of the amplification factor  $|\lambda|$  as a function of  $\omega_h^2 \Delta t_1^2 / 6$  for  $C_0 = 0.8$  and 1.02 in the revised algorithm.

to the direct-implicit algorithm [4], wherein the timestep constraint set by the stability of the highest frequency mode is completely relaxed, the timestep constraint for this semi-implicit orbit-averaged algorithm is only partially relaxed. Of course, there are residual timestep constraints on any algorithm set by accuracy considerations in calculating the particle trajectories and the dielectric response [8], e.g.,  $\mathbf{k} \cdot \mathbf{v} \Delta t < 1$ , so that an arbitrarily large timestep could never be used in practice. In Sections 3 and 4, we will describe how the  $\omega_h^2 \Delta t_i^2 \le 12$  constraint for the revised algorithm must compete with other timestep constraints that are sensitive to the specific physics application.

Our implementation of the semi-implicit orbit-averaged algorithm includes spatial filtering and spatial grid effects arising from the interpolations of the electric field force from the grid to the particles and the electron charge density from the particles to the grid. These alter the algorithm and the linear dispersion analysis in the same fashion as for the direct-implicit method [8, 9]. The spatial smoothing factors, e.g., a k-space filter like  $\exp(-k^2a^2)$ , and the interpolation factors that arise in  $\bar{n}_{N+1}^i - \langle n^e \rangle_{N+1}$  on the right side of Eqs. (3) and (4) should be matched by the same factors in the semi-implicit susceptibility. This is straightforward and we refer to [9] for explicit formulae. The dominant effect is a reduction of the effective value of  $\omega_h^2$  as  $(k \Delta x/2)^2$ or  $k^2a^2$  increases. As a consequence, the instability threshold on  $\omega_k^2 \Delta t_i^2$  becomes wavenumber dependent. Of course, thermal effects introduce additional dispersion; but this is physical. The Gaussian k-space filter can be suppressed by setting the smoothing factor a = 0. There is nothing in the dispersion analysis to indicate that this cannot be done, but we have not tried it. Having a tunable spatial filter has been useful in our experience with other particle codes, and we will make use of it here.

Before concluding this section, we comment on accuracy considerations associated with the semi-implicit susceptibility. A criterion for whether low-frequency and long wavelength modes of physical interest are being distorted can be derived from linear theory. The criterion derives from constraining the implicit susceptibility to be smaller than the linear electron susceptibility for the physical modes of interest. For example, consider the linear electron susceptibility for adiabatic electrons,  $\chi_e = 1/k^2 \lambda_e^2$ . The semi-

implicit susceptibility produces a small modification if  $C_0 k_{||}^2 \omega_{pe}^2 \Delta t_i^2 < 1/\lambda_e^2$ , which is equivalent to  $C_0 k_{||}^2 v_e^2 \Delta t_i^2 < 1$  and similar to the stricter stability condition Eq. (41) in Ref. [6] involving the largest wavenumber retained. Because the modes of interest must satisfy  $k_{||}^2 \Delta x_{||}^2 \le 1$  for good resolution, this constraint does not preclude  $\Delta t_e \le \Delta t_i$ . With a sufficiently small  $\Delta t_i$ , the longest, most important parallel wavelength modes will suffer an acceptably small distortion of the dielectric response because of the semi-implicit susceptibility.

#### 3. SIMULATION TEST CASES

The performance of the two-dimensional implementation of the revised semi-implicit orbit-averaged algorithm has been assessed in two types of test cases so far. The first test case was the unstable ITG mode investigated earlier in [7]. The second set of test cases was for a plasma that was stable to the ITG mode.

### 3.1. Ion-Temperature-Gradient Instability

In Figs. 5, 6, and 7, we show results from gyrokinetic simulations of an  $\eta_i = 4$  ITG instability, where  $\eta_i =$  $(d \ln T_i/dx)/(d \ln n_i/dx)$ . Simulation results from an explicit, electron subcycled gyrokinetic code are presented in Fig. 5. The semi-implicit orbit-averaged algorithm was used to produce the results shown in Figs. 6 and 7. With electron subcycling [10]. Poisson's equation and the electron equations of motion are solved together at each electron timestep  $\Delta t_e$ . The ions are advanced using a time-averaged electric field and a larger timestep that is an integer multiple of the electron timestep. The subcycling choreography and the stability properties differ slightly depending on whether the ratio of the ion timestep to the electron timestep is odd or even; odd is preferred [10]. The value of the subcycling parameter was  $N_* = \Delta t_i / \Delta t_e = 3$  for the simulation results shown in Figs. 5, 6, and 7. Because the ion gyrokinetic particle advance involves a much more complicated calculation than the electron drift-kinetic advance (the ion calculation is  $\sim 5$  times more time consuming), there is a great benefit in computational efficiency from electron subcycling it if is compatible with accuracy and stability considerations.

A  $16 \times 16$  grid was employed in all of the ITG test cases,  $T_e = T_i$ ,  $B_y/B_0 = 0.01$ , no magnetic shear, and  $\Delta x = \bar{\rho}_i$ , where  $\bar{\rho}_i = (T_i/m_i)^{1/2}/\Omega_i$  is the thermal ion Larmor radius. In addition, Lee's multiscale treatment of the ion temperature and density gradients was used [3]. There were 64 electrons and ions per cell used in the simulations shown in Figs. 5 and 6. There were 32 electrons and 64 ions per cell used in the orbit-averaged simulation whose results are displayed in Fig. 7. From linear theory, the complex frequency for the  $(k_x, k_y) = (1, -1)$  (in units of  $2\pi/L$ ) mode is

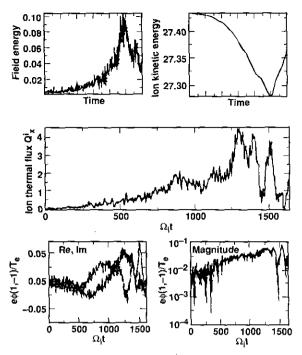


FIG. 5. Simulation results from an explicit subcycled simulation of the  $\eta_i=4$  instability in an unsheared magnetic field. The ion thermal flux has been summed over all the ions and normalized to the product of the number of ions times the ion temperature and the ion sound speed. The ion kinetic energy is the sum of all the ion kinetic energies in code units, and the field energy is the sum of the electric field and the ion polarization dielectric energy densities integrated over the volume in the same units as the ion kinetic energy. It is the nature of the ITG instability that the ion parallel kinetic energy is depleted at the expense of exciting the ITG modes to finite amplitudes. The real and imaginary parts and the magnitude of the Fourier amplitude of  $|e\phi/T_e|$  for the  $(k_x, k_y) = (1, -1)$  mode are also plotted as a function of time.

 $\omega/\Omega_i = 0.007 + 0.005i$  [7]. The unit charge per electron was twice that of the ions for the case with half as many electrons as ions. Further significant reduction of the number of orbit-averaged electrons relative to the number of ions did not lead to results with acceptable accuracy for the  $\eta_i = 4$  test case.

We note that there is reasonably good agreement between the simulation results shown in Figs. 5, 6, and 7. There is a significant degree of nonlinearity in the  $\eta_i = 4$  test case which influences the observed results and the best choice of simulation parameters. The mode frequency observed is significantly shifted by nonlinear effects from its linear value [7], and the maximum ion and electron velocities (derived from the  $\mathbf{E} \times \mathbf{B}$  velocity) are large enough to limit the choice of timesteps. In fact, we noted that there was sometimes a tendency for a strongly peaked potential structure to form which was accompanied by a strong velocity vortex for both species. Once a strong velocity vortex formed, smaller timesteps were needed in both the explicit and semi-implícit simulations to ensure that  $v \Delta t < \Delta x$  for both species. To the

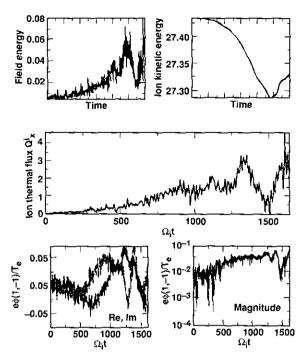


FIG. 6. Simulation results from a semi-implicit orbit-averaged simulation of the  $\eta_i = 4$  instability with equal numbers of electrons and ions.

extent that finite-amplitude effects narrow the disparity between the electron and ion velocities, so that the ions are not much slower than the electrons, the improvement in efficiency to be gained from subcycling and orbit averaging diminishes. This unstable  $\eta_i = 4$  example establishes that

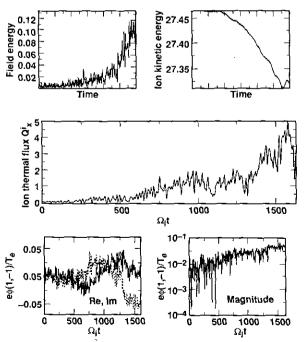


FIG. 7. Simulation results from a semi-implicit orbit-averaged simulation of the  $\eta_i = 4$  instability with half as many doubly-charged electrons as ions.

the revised semi-implicit algorithm can obtain results that are in satisfactory agreement with those of a published benchmark, but no improvement in computational efficiency was achieved in this particular case because of finite-amplitude effects.

### 3.2. Simulation of a Warm, Quiescent Plasma

The second test case set confirms some of the stability characteristics of the revised semi-implicit algorithm and demonstrates that a gain in computational efficiency can be obtained. In Fig. 8 we present simultion results for  $|e\phi/Te|$  for the  $(k_x, k_y) = (1, -1)$  mode as a function of time from three simulations with  $\eta_i = 1$ ,  $T_e = T_i$ ,  $\bar{\rho}_i = \Delta x$ ,  $16 \times 16$  grid, a smoothing factor  $\exp(-k^2a^2)$  with a = 1, and different choices of  $C_0$ ,  $N_*$ , and  $\omega_h \Delta t_i$ . In all three simulations there were 256 ions and 64 electrons per cell. In Fig. 8a,  $C_0 = 1.04$ ,  $N_* = 16$ , and  $\omega_h \Delta t_i = 1.86$ ; and the simulation is numerically stable. The simulation results in Fig. 8b reveal a numerical instability for  $C_0 = 0.5$ ,  $N_* = 16$ , and  $\omega_h \Delta t_i = 1.86$ ; and in Fig. 8c for  $C_0 = 1.04$ ,  $N_* = 64$ , and  $\omega_h \Delta t_i = 7.4$ . These results are consistent with the stability boundaries determined in Section 2.2.

We show results from simulations with no gradients in Fig. 9,  $T_e = T_i$ ,  $\bar{\rho}_i = \Delta x$ ,  $16 \times 16$  grid, 256 ions per cell, smoothing factor a = 1.7, and various choices for  $N_*$  and

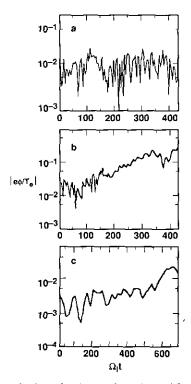
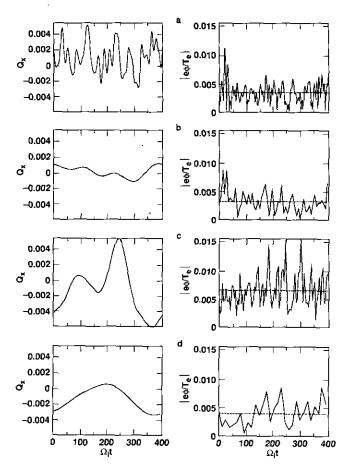


FIG. 8. Plots of  $|e\phi/T_e|$  for the  $(k_x, k_y) = (1, -1)$  Fourier mode from semi-implicit orbit-averaged simulations of a warm quiescent plasma for (a)  $C_0 = 1.04$  and  $\omega_h \Delta t_i = 1.86$ , (b)  $C_0 = 0.5$  and  $\omega_h \Delta t_i = 1.86$ , and (c)  $C_0 = 1.04$  and  $\omega_h \Delta t_i = 7.4$ .



**FIG. 9.** Simulation results for the normalized, time-filtered ion thermal flux  $Q_x$  in x and  $|e\phi/T_e|$  for the  $(k_x,k_y)=(1,-1)$  Fourier mode as functions of time for a warm quiescent plasma. (a) 256 electrons per cell,  $N_*=3$ , and the explicit subcycled algorithm. (b) 256 electrons per cell,  $N_*=25$ , and the semi-implicit orbit-averaged algorithm. (c) 64 electrons per cell,  $N_*=16$ , and the semi-implicit orbit-averaged algorithm. (d) 64 electrons per cell,  $N_*=50$ , and the semi-implicit orbit-averaged algorithm. The dashed curves on the  $|e\phi/T_e|$  plots are time averages.

the number of electrons per cell. These simulations of thermal noise are intended to show the influence of orbit averaging on important observables, such as the electric potential mode amplitudes and the ion thermal flux, and to address the questions of whether the orbit averaging reduces the noise and allows a reduction of the number of electrons. The results from one frame to the next should not be identical. We present results for  $|e\phi/T_e|$  for the  $(k_x, k_y) = (1, -1)$ mode and the spatially averaged ion thermal flux in the x direction normalized to the mean ion thermal energy times the sound speed as functions of time in Fig. 9. The deviations of the thermal flux from zero are due purely to statistical noise and thermal fluctuations. Both ions and electrons contribute to the statistical noise. The orbit averaging of the electrons can reduce only the electron contribution to the noise. For equal numbers of ions and electrons, increasing  $N_{\star}$  reduces the electron contribution to the statistical noise;

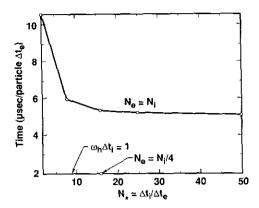


FIG. 10. Code timings (per particle per electron timestep) for semi-implicit orbit-averaged simulations of a warm quiescent plasma as a function of the subcycling parameter  $N_* = \Delta t_i/\Delta t_e$  for equivalent numbers of ions and electrons, and the effective code timing for one-fourth the number of electrons in a simulation that obtained similar physical results.

but the ion contribution to the noise should persist (Fig. 9). With  $N_* \gg 1$ , we were able to reduce the number of electrons significantly, e.g., from 256 to 64 per cell, while leaving the average thermal fluctuation levels in the same range (Figs. 9c and d). Furthermore, a numerically stable simulation was performed with  $\omega_h \Delta t_i = 5.8$  and  $N_* = 50$  for a smoothing factor a = 1.7.

Code timings on a single processor Cray 2 for the semi-implicit orbit-averaged algorithm with fully vectorized scatter [11] for the charge density accumulation are displayed in Fig. 10. The timings were calculated by dividing the processing time per electron timestep by the total number of particles  $(2N_i)$  for equal numbers of electrons and ions. The electron timestep is used so that a reference simulation with equal electron and ion timesteps and equal numbers of particles can be used, against which the code efficiency improvements due to electron subcycling and orbit-averaging can be measured. The reduction in code timing with increasing  $N_{*}$  is a result of the electron subcycling; i.e., we are not advancing the ions as often. The orbit averaging leads to an additional improvement in the total cost of the simulation by allowing the reduction of the number of electrons if  $N_{\star} \gg 1$  while maintaining an acceptable statistical noise level. The timing for advancing an individual electron or ion over its respective timestep remains the same. Thus, we have obtained an order of magnitude improvement in the code timing for a simulation of a warm quiescent plasma through the use of the semiimplicit orbit-averaged algorithm.

### 4. CONCLUDING REMARKS

In this paper, we have reported our experience in implementing a semi-implicit orbit-averaged gyrokinetic particle simulation algorithm. We found that our first algorithm introduced earlier in [2] was subject to a temporally growing error. This motivated the revised algorithm introduced and analyzed in Section 2. Simulation experience with the revised algorithm was reported in Section 3. The results of the simulation with the revised algorithm are in substantial agreement with the published results [7] and with those from a simulation with an explicit time-integration algorithm for the  $\eta_i = 4$  ITG instability. The simulation results reported in Section 3 also demonstrate good electron noise-reduction characteristics and a significant improvement in the code timings for orbit-averaged simulation of a warm quiescent plasma.

The algorithm implementation described here is restricted to the case of a possibly tilted magnetic field with no shear. Some comments on the possible extension of this algorithm to allow for magnetic shear are given in the Appendix. This work should be considered as a stepping stone in the development of more powerful particle simulation methods.

An important limitation on the applicability of the semi-implicit orbit-averaged gyrokinetic algorithm was discovered in the course of our work. The limitation derives from timestep constraints set by the physics of the applications. When the dominant timestep constraints are set by accuracy restrictions determined by equilibrium conditions (e.g., particle Courant conditions set by thermal velocities for the ions and electrons,  $v_i \Delta t_i < \Delta x_{||}$  and  $v_e \Delta t_e < \Delta x_{||}$ , where  $\Delta x_{||} = \Delta x/\theta$ ,  $v_{i,e}^2 = T_{i,e}/m_{i,e}$ , and  $\theta = B_y/B_0$  ( $\ll 1$ )), then there is a large disparity in timescales:  $\Delta t_i/\Delta t_e =$  $O(m_i/m_e)^{1/2}$  for  $T_i/T_e = O(1)$ . The orbit-averaged algorithm is designed to take advantage of this disparity to achieve an improvement in computational efficiency. Note that because we are specifically interested in physical circumstances for which ion Landau resonance might be important, viz.,  $\omega/k_{\parallel}v_i = O(1)$ , the timesteps for the evolution of the fields and the ion advance have been taken to be equal. If most of the electrons are moving much faster than the ions, we need to use a smaller timestep to accurately resolve their trajectories, but we can average their contributions to the charge and current densities in Maxwell's equations to reduce the computational costs. However, the simulation timesteps cannot be chosen without regard to finite-amplitude considerations. In particular, in the gyrokinetic simulation model the electron  $\mathbf{E} \times \mathbf{B}$  velocity  $V_E$ and the ion gyro-averaged  $\mathbf{E} \times \mathbf{B}$  velocity  $V_G$  must satisfy  $V_E \Delta t_e < \Delta x$  and  $V_G \Delta t_i < \Delta x$ . In general,  $V_G \leqslant V_E$ , but for  $k_{\perp}\bar{\rho}_{i} \leqslant 1$ ,  $V_{G}$  and  $V_{E}$  could be comparable; and there would be little disparity between the perturbed electron and ion drift velocities. This circumstance presents an impediment to the use of orbit averaging. For example, if  $(V_G/v_i)$  >  $(\Delta x/\Delta x_{\rm H})$ , then an ion timestep must be used to resolve the perturbed drift motion accurately that is smaller than that determined by resolving the equilibrium thermal streaming parallel to the magnetic field.

In the  $\eta_i=4$  ITG instability test case reported here, the wave amplitudes and the perturbed drift motion that were observed required the use of a relatively small timestep; and the value of  $N_*=\Delta t_i/\Delta t_e$  was relatively small, viz.,  $N_*=3$ . Relatively large electric potentials were observed:  $|e\phi/T_e| \le 0.18$ . The simulation timesteps were constrained by the Courant conditions on the peak velocity perturbations, which satisfied  $V_E \Delta t_e/\Delta x=0.3$  and  $V_G \Delta t_i/\Delta x=0.08$  for  $\Delta t_i/\Delta t_e=3$ . With these timesteps, the Courant conditions for the thermal velocities were easily satisfied,  $\theta v_e \Delta t_e/\Delta x=0.06$  and  $\theta v_i \Delta t_i/\Delta x=0.004$  for  $\theta=0.01$ . It should be noted that the simulations with the explicit and semi-implicit algorithms were both subject to these Courant conditions.

Simulations of the ITG instability modeling larger systems with weaker gradients than those used in our examples might be expected to saturate at smaller wave amplitudes and with weaker perturbed drifts on the basis of free-energy or mixing-length arguments [7]. Mixing-length arguments lead to estimates for the perturbed electric potential at saturation  $e\phi/T_e \sim 1/kL_n$  or  $1/kL_T$ , where k is the wavenumber appropriate to the fastest growing modes or the modes dominating the fluctuation spectrum. These arguments have been found useful in providing estimates, sometimes upper bounds, for fluid and particle simulations of turbulent transport. Because we know of no better a priori estimate of the saturated turbulence, we use the mixing-length estimate to express  $V_E/v_e = ck\phi/Bv_e \sim \rho_e/L_n$ , where  $\rho_e$  is the thermal electron Larmor radius. Thus,  $V_E/v_e$ is proportional to  $L_n^{-1}$ ; systems with weaker gradients might be more compatible with the use of large values of  $\Delta t_i$ and  $N_{+}$  in an orbit-averaged algorithm. In general, physics applications in which finite-amplitude effects are relatively weaker so that equilibrium-determined constraints are dominant and in which there is a large disparity between the ion timescales and the faster electron timescales will provide more fruitful opportunities for the use of the semi-implicit orbit-averaged algorithm.

# APPENDIX: EXTENSION OF THE ALGORITHM TO INCLUDE MAGNETIC SHEAR

The algorithms described in the text of this paper were restricted to having a tilted, unsheared magnetic field. A more relevant model of the magnetic field for physical applications is one in which there is magnetic shear,  $B_y = B_0(x - x_0)/L_s$ , where  $L_s$  is the magnetic shear length. The methodology of the particle pushing and charge accumulation remains unchanged from the algorithm described in the text, but the solution of the field equation must be revised for magnetic shear. In the presence of magnetic shear, the explicit dependence of  $\nabla_{||}$  on x,  $\nabla_{||} = (\partial/\partial z) + (x - x_0) L_s^{-1}(\partial/\partial y)$ , frustrates the use of Fourier transforms in x to solve the semi-implicit Poisson equation,

Eq. (4). Here we suggest a method for solving Eq. (4) when there is magnetic shear.

If one cannot Fourier transform in x, then the evaluation of  $-(\omega_{pi}^2/\Omega_i^2)$   $\nabla_{\perp}^2 g = \lambda_i^{-2} [1 - I_0 \exp(-b)]$ , where  $\lambda_i = v_i/\omega_{pi}$ , cannot be performed in  $\mathbf{k}$  space. Research on gyrofluid models [12] has demonstrated the utility of the Padé approximation,  $I_0 \exp(-b) \approx (1+b^2)^{-1}$  for  $0 \le b \le 9$ . For simplicity we drop the vacuum polarization term  $\nabla^2$  in Eq. (3) (which is typically small), use the Padé approximation for  $I_0 \exp(-b)$ , and apply the operator  $\lambda_i^2 (1+b) = \lambda_i^2 (1-\bar{\rho}_i \nabla_{\perp}^2)$  to both sides of Eq. (3) to obtain

$$\begin{aligned} \left\{ -\bar{\rho}_i \nabla_{\perp}^2 + \lambda_i^2 (1 - \bar{\rho}_i \nabla_{\perp}^2) \nabla \cdot \left[ C_0 \chi_e \cdot \nabla \right] \right\} \tilde{\phi}_{N+1} \\ &= \lambda_i^2 (1 - \bar{\rho}_i \nabla_{\perp}^2) S, \end{aligned} \tag{A1}$$

where  $S = 4\pi e(\bar{n}_{N+1}^i - \langle n^e \rangle_{N+1}) - \nabla \cdot [C_0 \chi_e \cdot \nabla \phi_N]$ , which has explicit dependence on x through  $\nabla_{||}$  (see Eq. (4)). The expression in the curly bracket on the left side of Eq. (A1) is an elliptic operator whose finite-difference representation in x leads to a matrix that is sparse and banded. Fourier transforms can still be used to evaluate y and z derivatives in k space. We expect that the solution of the resulting sparsematrix equation will be amenable to one of the many methods that exist for such equations, but we have made no attempt to implement this algorithm extension as yet.

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